

# LHC black holes and the end of high- $p_T$ physics

- How could this be possible?
- Why is it the end of high- $p_T$  physics?
- Simulations
- Results
- Uncertainties
- Conclusion

Work done with Leif Lönnblad and Torsten Åkesson

Lund 2005.05.13 Malin Sjödahl

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- In nature there is a hierarchy between the EW scale  $M_{EW} \sim 10^2$  GeV and the Planck scale  $\sim 10^{19}$  GeV
- This could be solved if gravity was allowed to propagate in a lager space than the other forces



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- Modified version by Randall and Sundrum (RS) with curved space and only one extra dimension





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- This means that energetic partons which otherwise would have undergone a QCD scattering may be trapped behind the event horizon
- But the partons need to be well localized also in transverse direction
- Hence the end of high  $p_T$ -physics!



#### **Simulations:**



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• **PYTHIA** - favorite event generator

Creates QCD events, but we cut some away according to

$$\frac{d\sigma_{\rm QCD}(Q^2)}{dQ^2 d\hat{s}} = \int dx_1 dx_2 \sum_{i,j} f_i(x_1, Q^2) f_j(x_2, Q^2) \frac{\hat{\sigma}_{ij}^{\rm QCD}(\hat{s}, Q^2)}{dQ^2} \times \delta(\hat{s} - x_1 x_2 s) \left[ 1 - \Theta(\sqrt{\hat{s}} - M_{\rm min})\Theta(Q^2 - \frac{const}{r_{Sch}^2(\hat{s})}) \right]$$





Creates a black hole out of two partons according to

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Adjusts temperature according to mass,  $T \propto 1/r(m)$ 

















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- When does a black hole form?
  - $\lambda \lesssim r_{Sch}$

Making this condition stronger gives fewer particles in the low energy end

- $\rightarrow$  Doesn't change chance to observe QCD drop much
- How do we terminate the decay of the black hole?
  - 2 particles when we reach  $M_p$
  - More particles  $\rightarrow$  Fewer energetic particles
  - $\rightarrow$  Better chance to observe QCD drop



- Gravity important before black hole production? Yes, we should have non black hole gravitational events.
  - $\rightarrow$  Affect the low transverse momentum end of spectrum.
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• More





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- But this may (naively) be camouflaged by the black hole decay products for rather large transverse energies
- More sophisticated methods are needed to observe QCD drop
- If the black hole is stable on collider timescales (RS) there is no radiation to camouflage the drop

